

Lecture 35: Radiative Lifetime Ref. Chapter 10.1 & 10.2

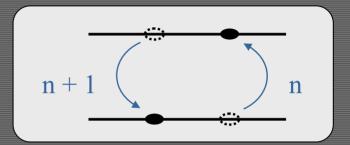




Introduction

 Ordinarily if we use the two levels for our basis functions, then the Hamiltonian will be diagonal. That would mean no transitions, because levels 1 and 2 would be completely decoupled from each other. For transition between levels, you need off diagonal elements such that there is coupling between levels. So the postulation is that because of the electromagnetic noise that is present, electron will be perturbed which is the scattering potential in [H]. The problem is that if you put any real potential, then its complex conjugate will be the other off-diagonal element; This means that the transitions form 1 to 2 and from 2 to 1 are equal. How ever what happens in equilibrium is that electrons relax at a much higher rate than they get excited.

Emission and Absorption



$$\begin{bmatrix} i\hbar \begin{Bmatrix} \psi_1 \\ \psi_2 \end{bmatrix} = \begin{bmatrix} \varepsilon_1 & \begin{bmatrix} U_S \end{bmatrix}_{12} \\ \begin{bmatrix} U_S \end{bmatrix}_{21} & \varepsilon_2 \end{bmatrix} \begin{Bmatrix} \psi_1 \\ \psi_2 \end{bmatrix}$$

• The number of upwards transitions is proportional to n and the number of downward transitions to n+1, where

$$n = \frac{1}{e^{\hbar\omega/k_BT} + 1}$$
 at T=0, n=0

Multi-Particle Picture

• That said, the emission/absorption process is much easier to understand in terms of a multi-particle electron-photon picture (as discussed in the previous lecture)

Multi-Particle Levels

N-1 photon

N photon

N+1 photon

$$\underline{\varepsilon_2 + (N)\hbar\omega} \quad \underline{\varepsilon_1 + (N+1)\hbar\omega}$$

$$\frac{\varepsilon_2 + (N-1)\hbar\omega}{\varepsilon_1 + (N-1)\hbar\omega} \quad \frac{\varepsilon_1 + (N)\hbar\omega}{\varepsilon_1 + (N-1)\hbar\omega}$$

$$\varepsilon_1 + (N-1)\hbar a$$

• If you think of the atom as being in one big box, then photons in the box can be described by plane waves:
$$\cos(\vec{\beta} \cdot \vec{r} - \omega(\vec{\beta})t)$$

ullet The relationship between ω and β is linear. We have:

$$\hbar \omega (\vec{\beta}) = \hbar c \beta$$

$$slope = c$$

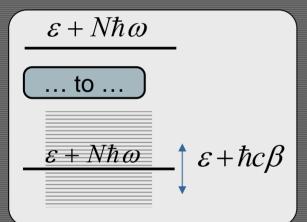
 One photon state means having one photon at a particular frequency. N photon state means having N photons at that frequency.

Note: With photons there is no exclusion principle

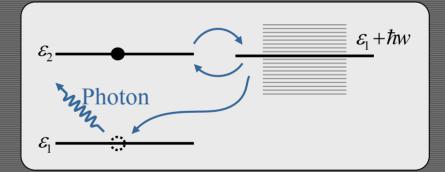
Dissipation

• The continuous distribution of photons alters the 1-level picture from...

• Looking at the 0-photon state, when we put an electron in the upper level, ε_2 , it likes to spread out (broaden) into the adjacent level and then fall back down to ε_1 giving up a photon in the process



Dissipation into Photon Reservoir

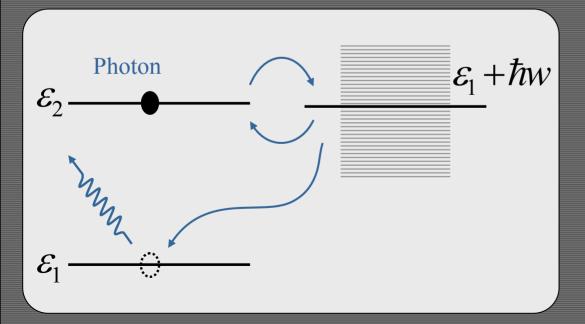


- This has important implications for photon emission from excited states
- Importantly, this behavior can be viewed much the same as dissipation into a contact
- So the question is how one calculates the rate at which the electron goes away when it's put in a state.

Lifetime

• Because adjacent levels act as photon reservoirs we may use the Green's function formalism to derive the radiative lifetime of an electron

Dissipation into Photon Reservoir



 So we can use the selfenergy expression

$$\Sigma = \tau g \tau^+$$

and to get an estimate of the lifetime take

$$\hbar/\tau_r = \Gamma_1 = \tau a \tau^+$$

The process may be viewed as $|2,0ph> \rightarrow |1,1ph>$ where radiative lifetime is like another self-energy

Coupling

- Note: τ is a physical quantity which constitutes the coupling between levels and τ_{γ} is the escape rate of electron from one state to another.
- Formally, therefore the broadening of the 0-photon state is given as: $\Gamma_1 = \tau a \, \tau^+ = \sum_{\bar{\beta}} \left| \tau \left(\bar{\beta} \right) \right|^2 2 \pi \delta \left(E \varepsilon_1 \hbar \, \omega(\bar{\beta}) \right)$
- Where $2\pi\delta(E-\varepsilon_1-\hbar w(\vec{\beta}))$ is the photon reservoir spectral function a.
- Imposing periodic boundary this summation may be transformed into an integral

$$\begin{split} \hbar/\tau_r &= \int_0^\infty \int_0^\pi \int_0^{2\pi} \left[\frac{\beta^2 d\beta \sin\theta d\theta d\phi}{8\pi^3} V \Big| \tau \Big(\vec{\beta} \Big)^2 \times 2\pi \delta \Big(\varepsilon_2 - \varepsilon_1 - \hbar \omega (\vec{\beta}) \Big) \right] \\ &= \int_0^\infty \int_0^\pi \int_0^{2\pi} \left[\frac{\omega^2 d\omega \sin\theta d\theta d\phi}{8\pi^3 c^3} V \Big| \tau \Big(\vec{\beta} \Big)^2 \times 2\pi \delta \Big(\varepsilon_2 - \varepsilon_1 - \hbar \omega (\vec{\beta}) \Big) \right] \\ &= \int_0^\pi \int_0^{2\pi} \frac{\sin\theta d\theta d\phi}{8\pi^3 c^3} \Big| \tau \Big(\vec{\beta} \Big)^2 \Big(\frac{\varepsilon_2 - \varepsilon_1}{\hbar} \Big)^2 \times \frac{2\pi}{\hbar} V \qquad \text{(V is the volume)} \end{split}$$

Deriving the Coupling

- To arrive at an expression for coupling between the 0-photon and 1-photon state we must consider the potential that an electron feels due to one photon
- The electric field of a single photon with wave vector $\vec{\beta}$ is

$$\vec{\varepsilon} = \hat{v}\varepsilon_0 \sin\left(\vec{\beta} \cdot \vec{r} - \omega(\vec{\beta})t\right)$$

We want to know if the box had just one photon, then what electric field would be present. One way to do this is to look at the total energy in the EM wave:

$$\varepsilon_{Total} = \varepsilon_E^2 + \varepsilon_M^2 = \frac{V}{2} \varepsilon \varepsilon_0^2 = \hbar \omega$$

where V is the volume of the box containing the photon

• Thus,
$$\left| \varepsilon_0 \right| = \sqrt{\frac{2\hbar\omega}{\varepsilon V}}$$

 Now, to write the potential we use the well known relationship

$$\vec{E} = \frac{-\partial \vec{A}}{\partial t}$$

$$\therefore \vec{A} = \hat{v}A_0 \cos(\vec{\beta} \cdot \vec{r} - \omega(\vec{\beta})t)$$

$$\left| A_0 \right| = \sqrt{\frac{2\hbar}{\omega\varepsilon V}}$$

Vector Potentials and the Schrödinger Equation

- Note, however, that the potential created by a photon is a vector potential and may not be simply added to the Schrödinger equation as the typical electro-static scalar potential we are familiar with.
- General Approach:
- → A scalar potential, U, is included in the Hamiltonian as

$$H = \frac{p^2}{2m} + U$$

 \rightarrow and the EM vector potential, \bar{A} , as

$$H = \frac{\left(\vec{p} + q\vec{A}\right)^2}{2m} \quad \boxed{1}$$

 To find the coupling term we need to know how much the Hamiltonian changed due to addition of one photon. Expanding 1:

$$H = \frac{q}{2m} (\vec{A} \cdot \vec{p} + \vec{p} \cdot \vec{A}) + \frac{q^2}{2m} \vec{A} \cdot \vec{A}$$

 $\vec{A} \cdot \vec{A} \approx 0$ Negligible

$$\vec{p} \cdot \vec{A} = \vec{A} \cdot \vec{p}$$
 if $\vec{\nabla} \cdot \vec{A} = 0$

• So the Hamiltonian becomes:

$$H = \frac{q}{m} \left(\vec{A} \cdot \vec{p} \right)$$

Coupling Element

• As done earlier, we adopt

$$\vec{A} = \hat{v}A_0 \cos\left(\vec{\beta} \cdot \vec{r} - \omega(\vec{\beta})t\right)$$

thus the addition to Hamiltonian is

$$H = \frac{q}{m} \left(\vec{A} \cdot \vec{p} \right) = qA_0 / 2m \left(\vec{p} \cdot \hat{v} \right) \left(e^{+i\vec{\beta} \cdot \vec{r}} e^{-i\omega(\beta)t} + e^{-i\vec{\beta} \cdot \vec{r}} e^{i\omega(\beta)t} \right)$$

• Furthermore we denote $e^{+i\bar{\beta}\cdot\bar{r}}e^{-i\omega(\beta)t}$ as the absorption term and $e^{-i\bar{\beta}\cdot\bar{r}}e^{i\omega(\beta)t}$ as the emission term.

$$i\hbar \begin{cases} \psi_1 \\ \psi_2 \end{cases} = \begin{bmatrix} \varepsilon_1 & [U_S]_{12} \\ [U_S]_{21} & \varepsilon_2 \end{bmatrix} \begin{cases} \psi_1 \\ \psi_2 \end{cases}$$
 [H]

• So for the coupling element we have: $H_{12} = \frac{qA_0}{2m} \int dr^3 \phi_1^* \left(\vec{p} \cdot \hat{v} e^{\pm i\vec{\beta} \cdot \vec{r}} \right) \phi_2$