

ECE695: Reliability Physics of Nano-Transistors
Lecture 37: Effect of Radiation Induced Charge
-- Soft Errors

Muhammad Ashraful Alam
alam@purdue.edu

Copyright 2013

This material is copyrighted by M. Alam under the following Creative Commons license:



Attribution-NonCommercial-ShareAlike 2.5 Generic (CC BY-NC-SA 2.5)

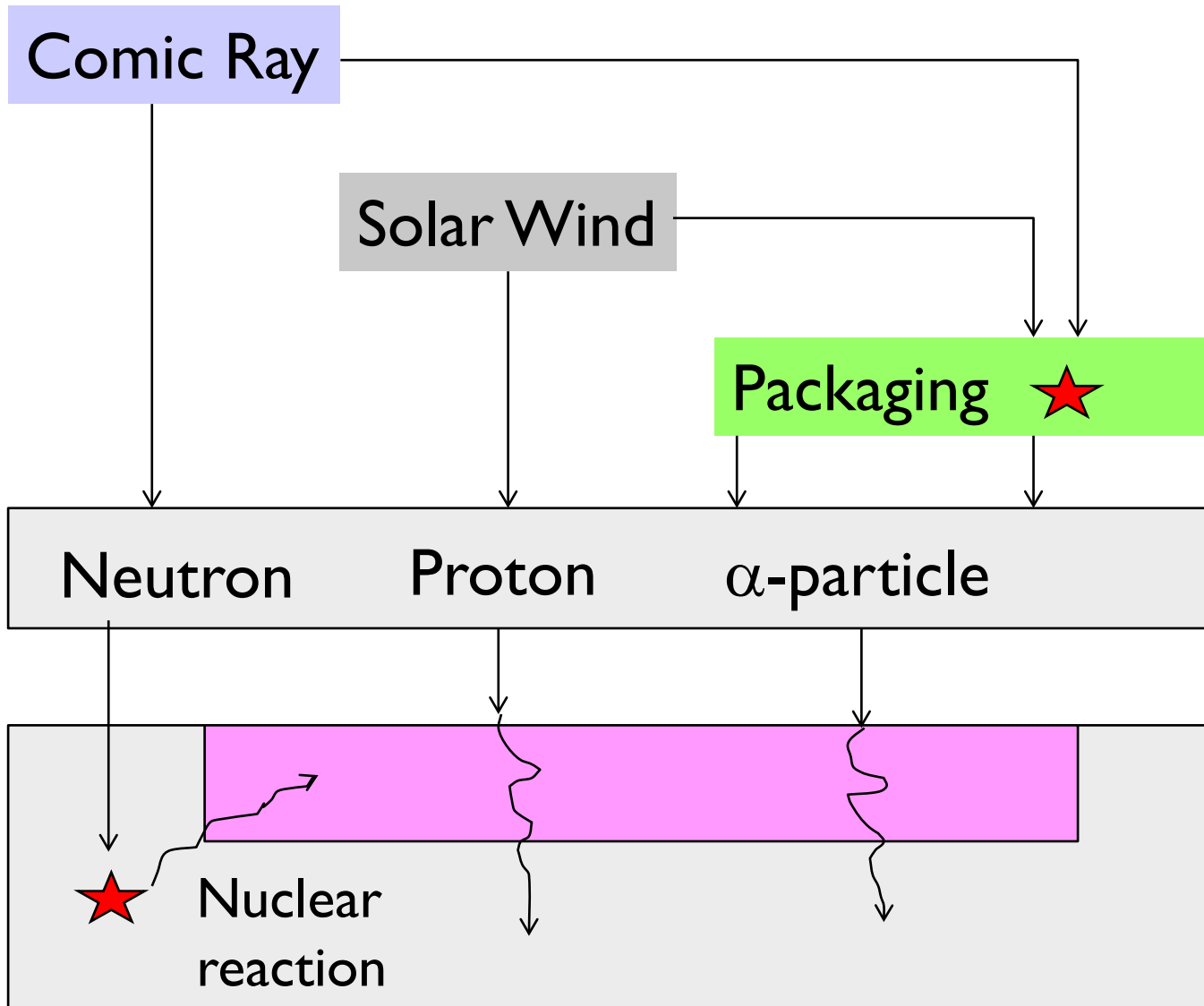
Conditions for using these materials is described at

<http://creativecommons.org/licenses/by-nc-sa/2.5/>

Outline

1. Sources of radiation
2. Basics of charge generation and perturbation of potential
3. Charges and junctions
4. Conclusions

Outline



Sources of Radiation

Cosmic Ray



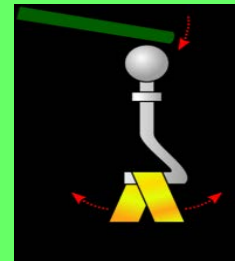
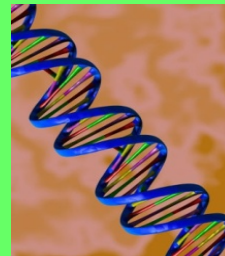
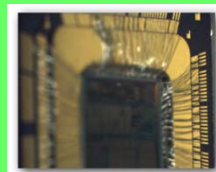
Solar Wind

- 4/17 Satellite yrs
- Telestar I < 1yr
- 1 fail/flight
- Exponential with altitude



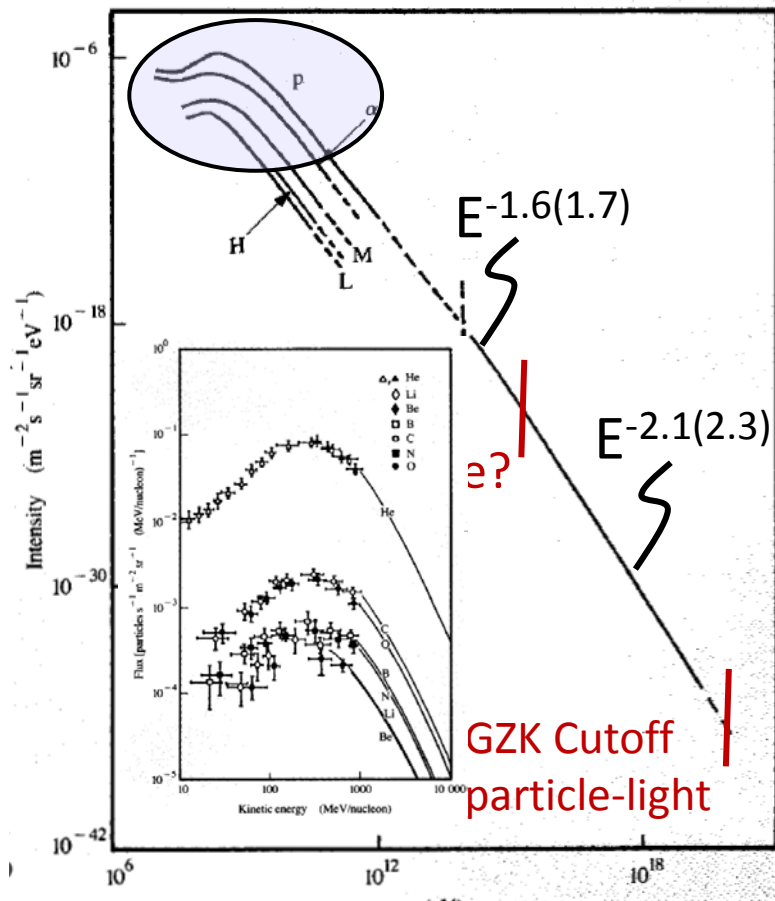
Earth Sources

- mutation
- foil electrometer



Source 1: Cosmic Ray (above atmosphere)

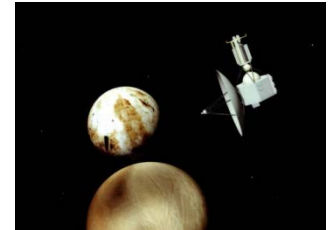
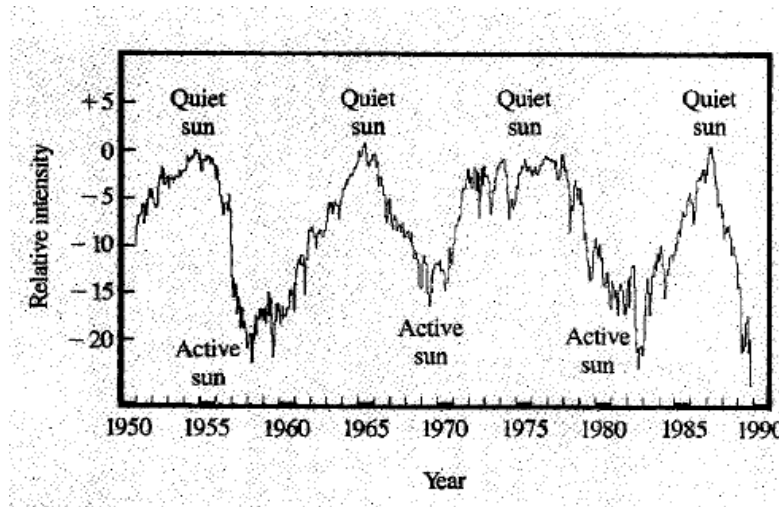
- 92% proton, 6% alpha (or, 70% proton, 30% neutron)
- 2% pion, muon, positron, Li, B, e, B, Pb, Fe(1% to earth, physics testbed)
- $1/\text{cm}^2$ integrated above 1 GeV, inconsequential



Isotropic:

- $\rho \times d = 4.2 \text{ g/cm}^2$
- ρ (space; C, N, O) $\sim 1.7 \times 10^{-26} \text{ g/cm}^3$
- Distance = $c \times \text{time}$
- Scattering time ~ 300 million-years (equivalent from Lab Expt).
- Isotropy indicates vast distances!

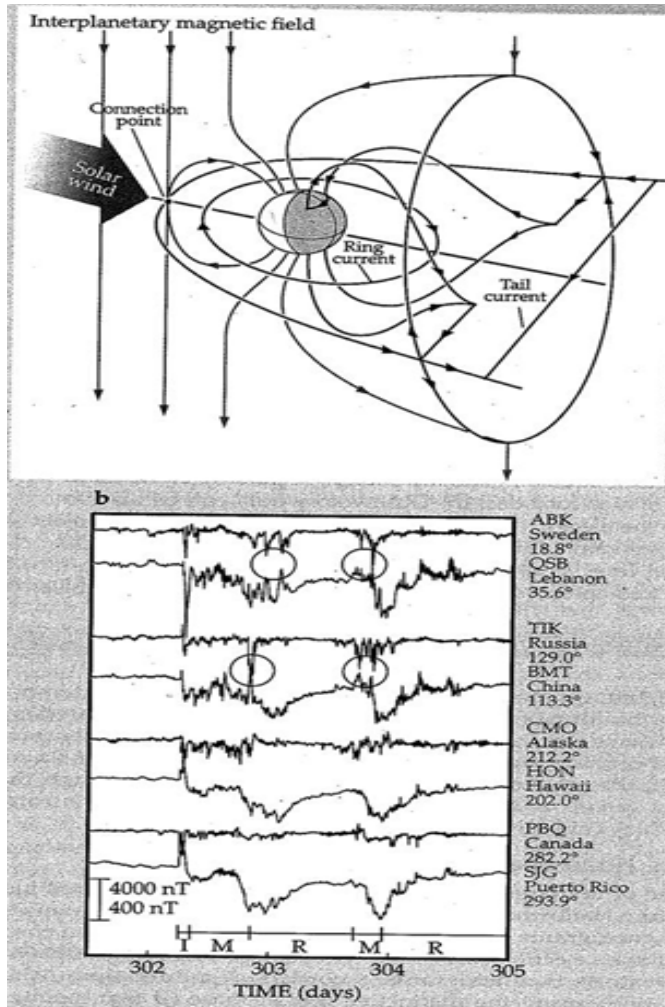
Sources 2: Solar Wind



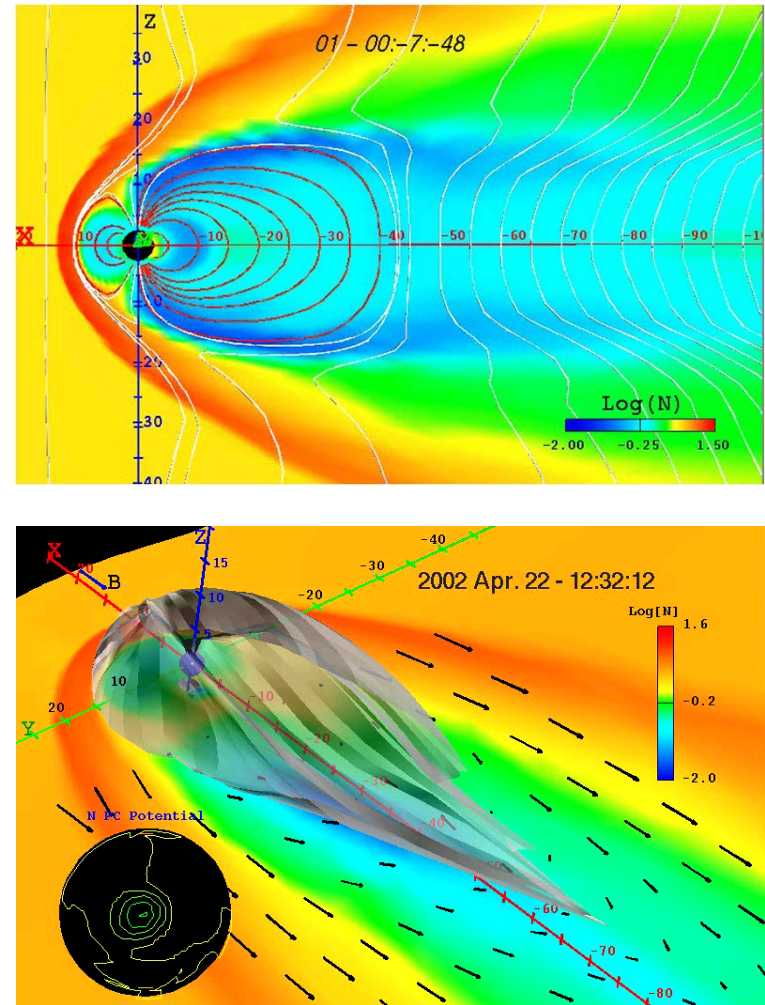
- 90B one Mton bombs; powers oceans, photosynthesis, atmosphere, etc.
- 2000 – 1 millions/cm²-sec < 1 GeV (anisotropic)
- 1859: burned telegraph wire, forest fire;
- 1989: Blackout in North-Eastern America; 11k/19k Satellites lost contact
- 1365 W/cm² at surface, 450 reaches earth, 300 reaches surface. Particles reach earth in 4 days

Asymmetry of Solar Wind

Measurement

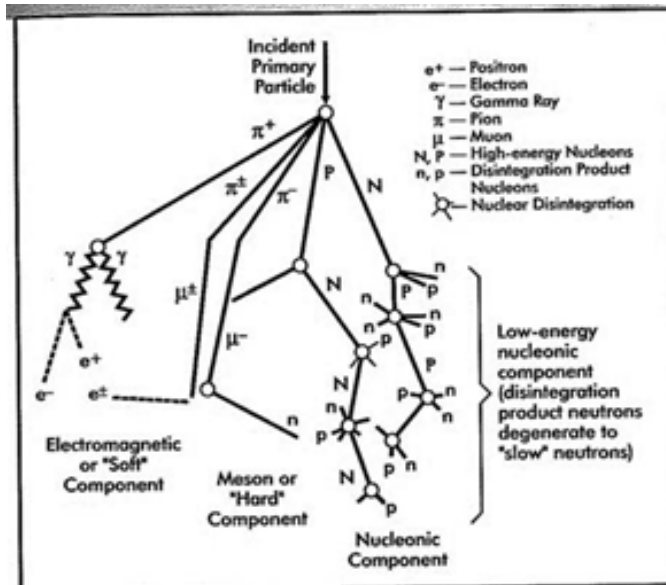
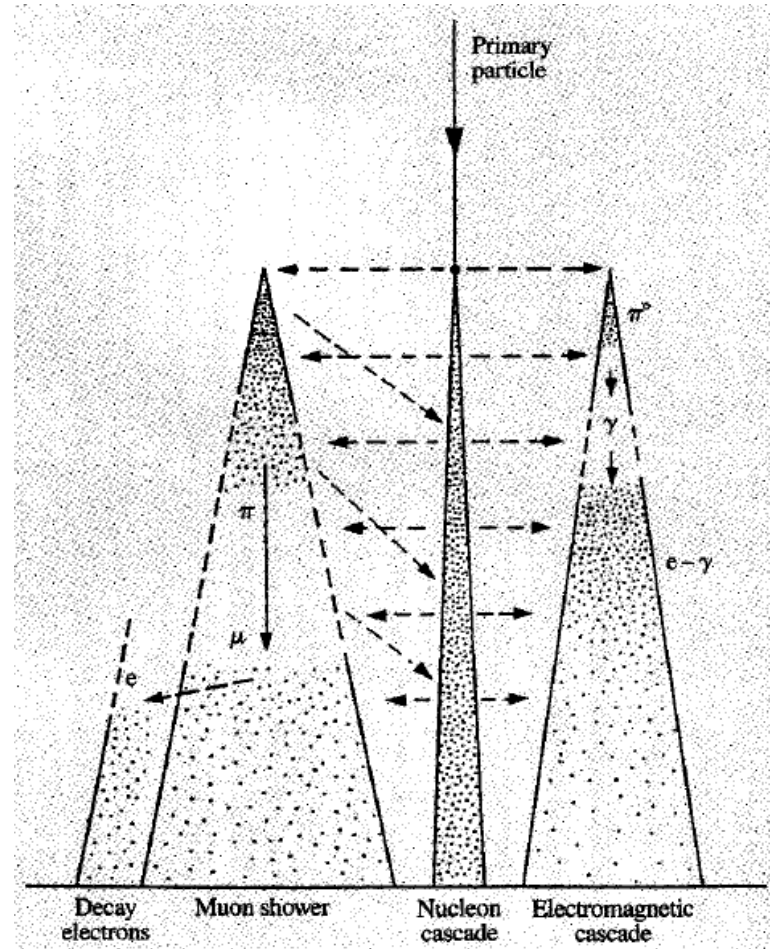
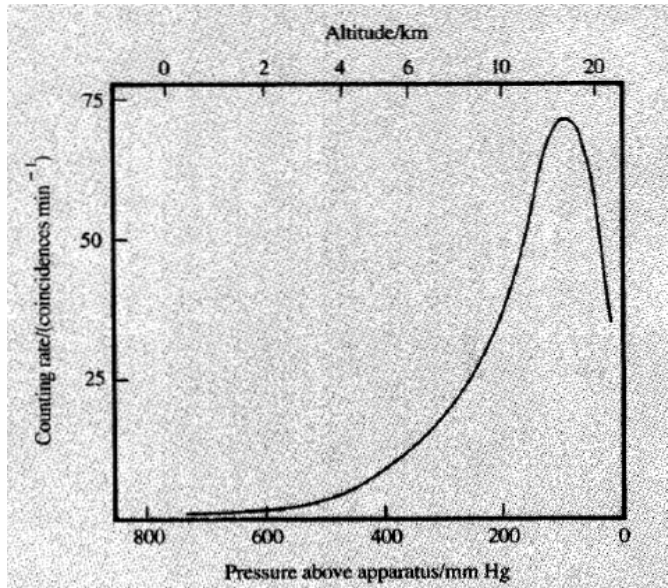


Simulation



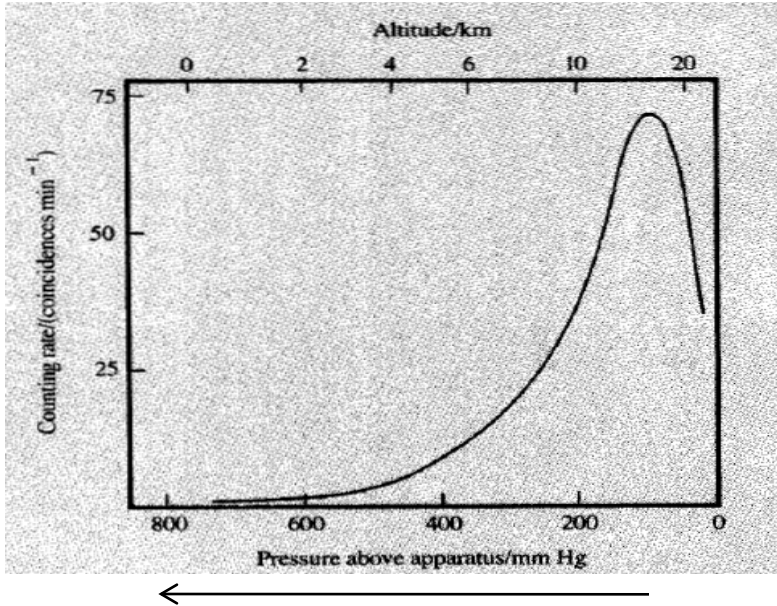
http://spp.astro.umd.edu/Rb_education/index.htm

Altitude Dependent Radiation Flux



Original – Hess Nobel Prize
Full curve – Pfitzner curve

Altitude Dependent Radiation Flux



Empirical Model:

$$I_2 = I_1 \exp\left(\frac{A_1(@ H_1 - A_2 @ H_2)}{L}\right)$$

$$A = 1033 - (0.03648H) + (4.26 \times 10^{-7} H^2) \text{ g/cm}^2$$

$$L_e = 100 \text{ g/cm}^2, L_p = 110 \text{ g/cm}^2$$

$$L_n = 148 \text{ g/cm}^2, L_\mu = 520 \text{ g/cm}^2$$

Brian solid slab model:

Atomic weight 14.48

Atomic number 7.22

Ionization 92.8 eV – reminiscent of Nitrogen

Bethe formula for energy loss:
 $dE/dx \sim \ln(E)/E$ (see Lec. 38)

Example 1: *Denver*, $H = 5280 \text{ ft}$

$$A_1(H = 0) = 1033$$

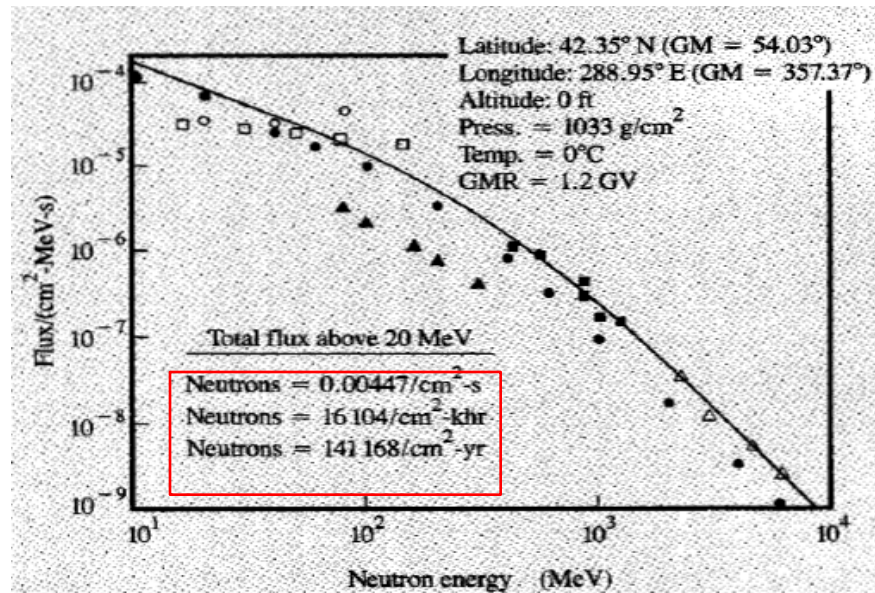
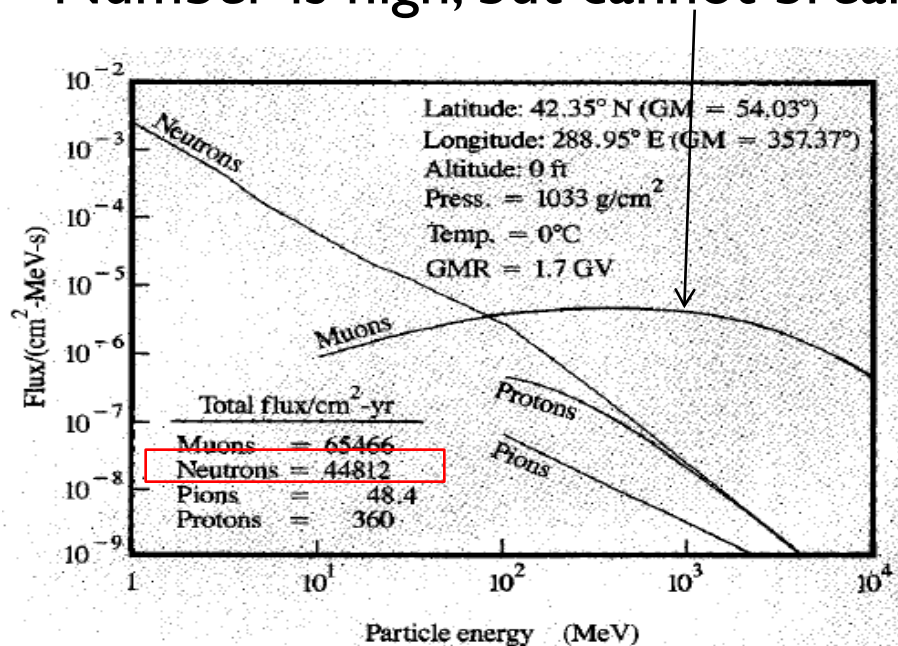
$$A_2(H = 5280) = 862$$

$$I @ 5280 \text{ ft} = \exp(1033 - 862) / 148 = 3.4$$

Example 2: At 1.5km (Airlines): 100x increase

Radiation Intensity at the Earth Level

Number is high, but cannot break anything (LET, dE/dx small)



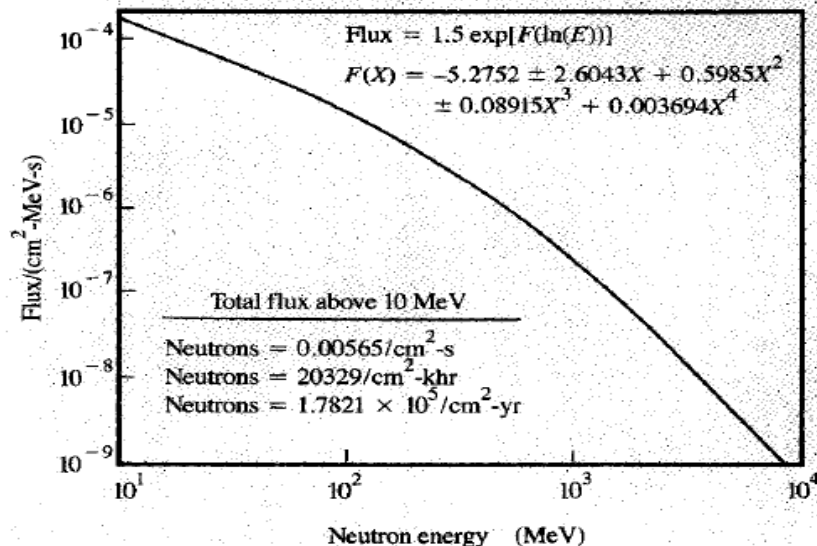
Empirical Model (Energy Distribution)

$$I_2 = I_1 \exp\left(\frac{A_1 - A_2}{L}\right) \times 1.5 \exp F(x) \quad x \equiv \ln E$$

$$F(x) \equiv -5.28 \pm 2.6x + 0.6x^2 \pm 0.09x^3 + 0.00369x^4$$

$$I_n (> 20 \text{ MeV}) = 0.5 - 1.5 \times 10^5 \text{ cm}^{-2} \text{ yr}^{-1}$$

$$I_\mu \gg I_n \quad \text{but} \quad (dE/dx)_\mu > (dE/dx)_n$$



Sources of Radiation 3: Packaging

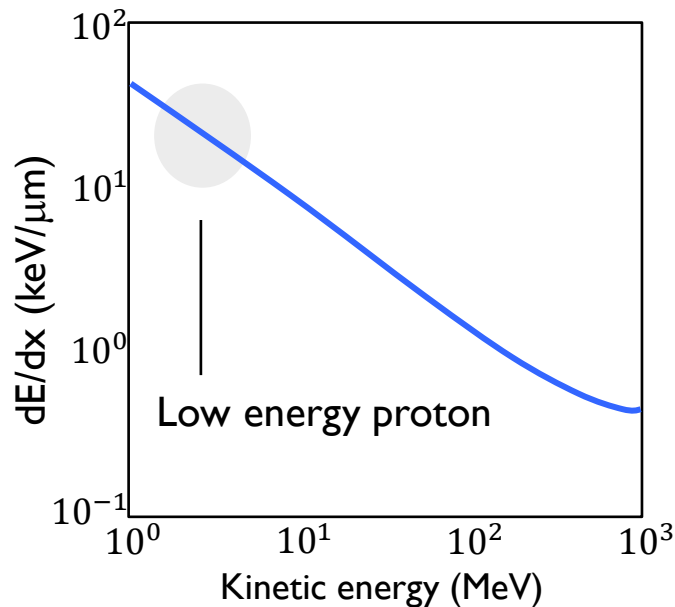
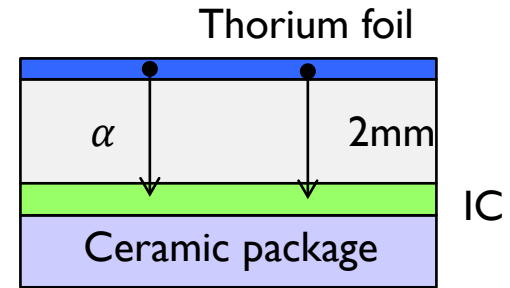
Earth Sources



Thermal Neutron in
BPSG



Proton on Li –
Thermal Neutron



Alpha particle fluxes (#/cm²-hr)

Processed Wafers	0.0009
Cu Metal (thick)	0.0019
Al Metal (thick)	0.0014
Mold compound	0.024-0.002
Underfill	0.002-0.0009
Pb-solders	7.2 -0.002
Ceramic package	0.0011

Slow particles from IC are devastating, previously did not survive

Sources of Radiation: Packaging

Soft Errors in Advanced Semiconductor Devices—Part I: The Three Radiation Sources

Robert C. Baumann, *Senior Member, IEEE*

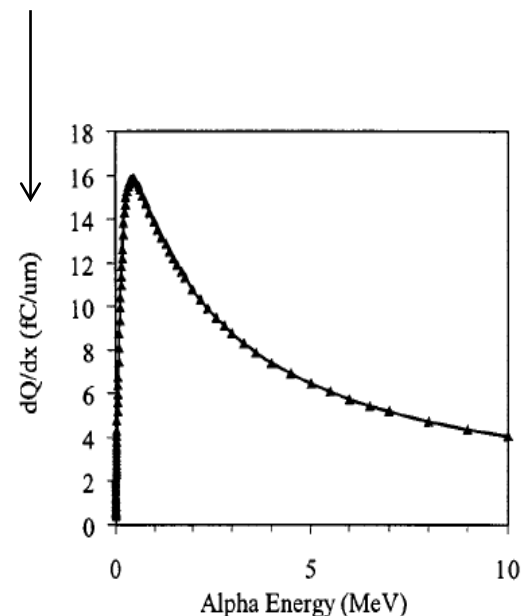
TABLE I
THE URANIUM SERIES

Species	Half-life	Mode	Energy (MeV)
U-238	4.47x10⁹ yrs	α	4.196(77), 4.149(23)
Th-234	24.1 days	β	
Pa-234	6.69 hrs	β	
U-234	2.45x10 ⁵ yrs	α	4.774(72), 4.723(28)
Th-230	7.54x10 ⁴ yrs	α	4.688(74), 4.621(26)
Ra-226	1.60x10 ³ yrs	α	4.785(95), 4.602(5)
Rn-222	3.82 days	α	5.490
Po-218	3.05 min	α	6.002
Pb-214	26.8 min	β	
Bi-214	19.7 min	β	
Po-214	164 usec	α	7.687
Pb-210	22.3 yrs	β	
Bi-210	5.01 days	β	
Po-210	138.4 days	α	5.305
Pb-206	Stable		

TABLE II
THE THORIUM SERIES

Species	Half-life	Mode	Energy (MeV)
Th-232	1.41x10¹⁰ yrs	α	4.016(77), 3.957(23)
Ra-228	5.76 yrs	β	
Ac-228	6.13 hrs	β	
Th-228	1.91 yrs	α	5.426(71), 5.343(29)
Ra-224	3.66 days	α	5.686(94), 5.449(6)
Rn-220	55.6 sec	α	6.288
Po-216	0.15 sec	α	6.779
Pb-212	10.64 hrs	β	
Bi-212	60.60 min	β (64)	
Bi-212	2.251	α(36)	6.336(57), 6.297(43)
		V	
Po-212	0.30 usec	α	8.785
Pb-208	stable		
		V	
Tl-208	3.05 min	β	
Pb-208	stable		

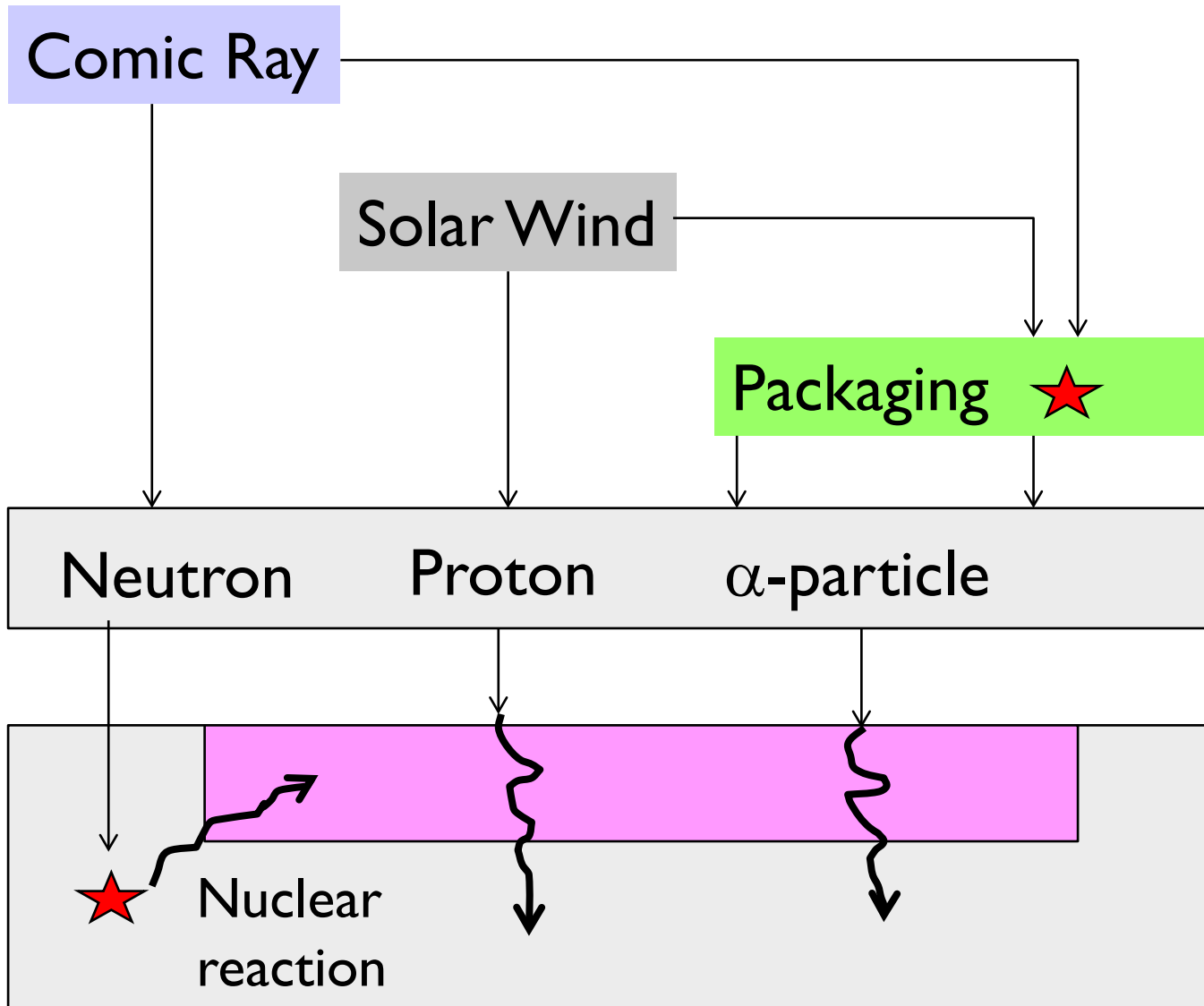
Charge generated



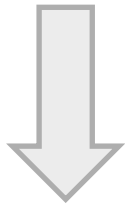
Outline

1. Sources of radiation
2. Charge generation and potential fluctuation
3. Junctions and critical charges
4. Conclusions

Outline

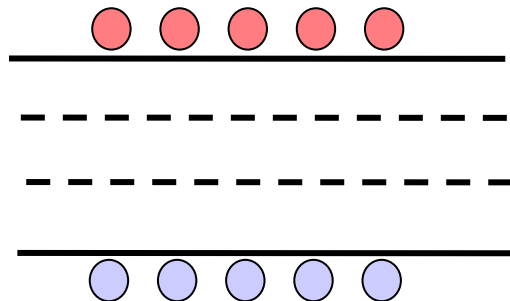
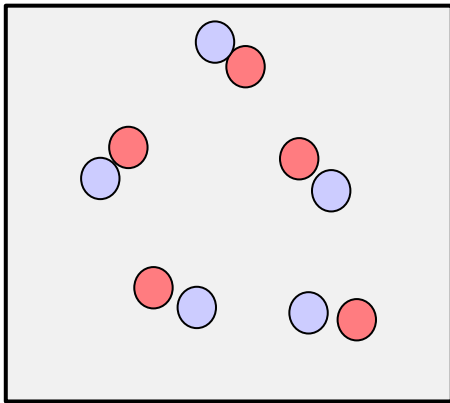


Generation of Electron-Hole Pairs



$$\varepsilon_i(\omega) = \frac{q^2 P_0^2 \sqrt{\mu(\hbar\omega - E_g)}}{\pi m_0^2 \omega^2 \hbar^3}$$

$$P_0^2 = \frac{m_0 E_g}{2} \left(\frac{m_0}{m_c} - 1 \right) \frac{3E_g + 3\Delta}{3E_g + 2\Delta}$$



$$n_{ph} = \frac{I_0 (e^{-\alpha x} - 1)}{c'} = n = p$$

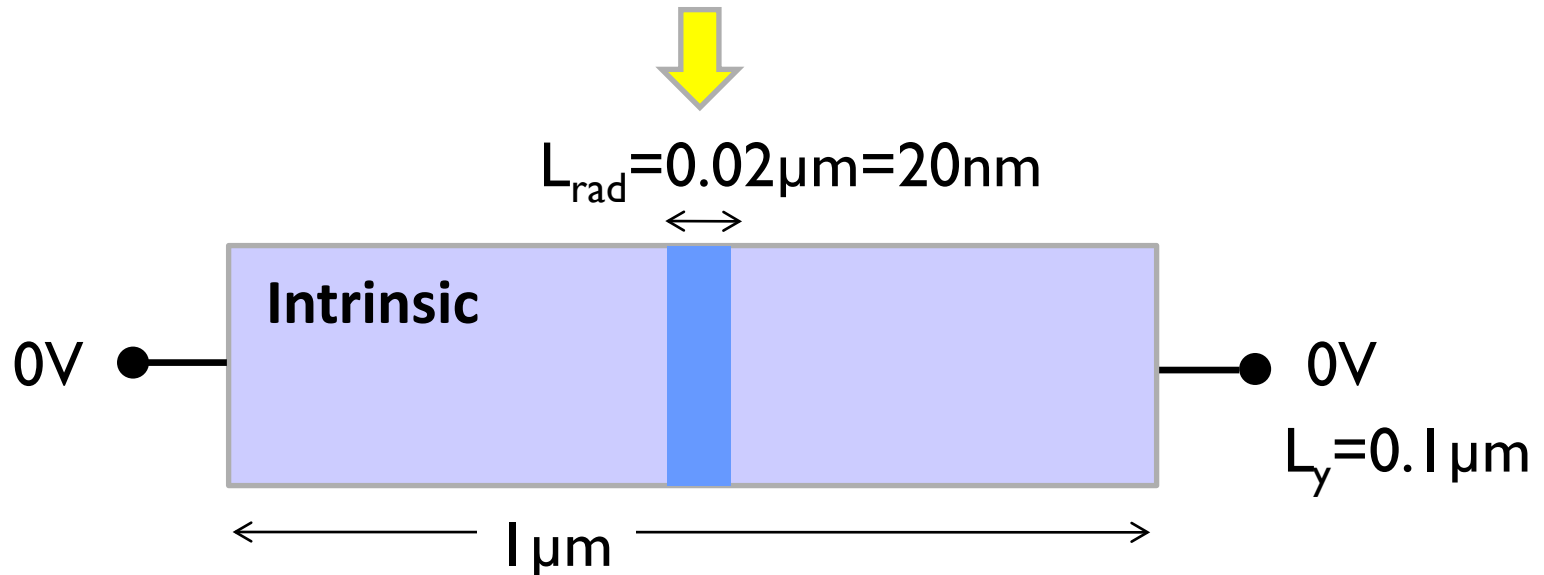
$$\alpha_{dir} = \frac{q^2 (2\mu)^{3/2} \sqrt{(\hbar\omega - E_g)}}{nch^2 m_e}$$

$$= 2 \times 10^4 \sqrt{(\hbar\omega - E_g)} \text{ cm}^{-1}$$

$$\alpha_{indir} = \frac{4 q^2 \mu^{5/2} (\hbar\omega - E_g)^{3/2}}{3 nch^2 m_e m_h hv}$$

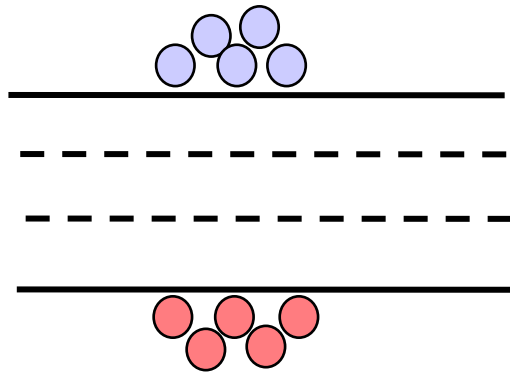
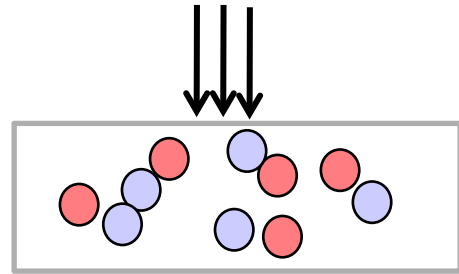
$$= 1.3 \times 10^4 \frac{(\hbar\omega - E_g)^{3/2}}{hv} \text{ cm}^{-1}$$

Intrinsic structure: particle strike



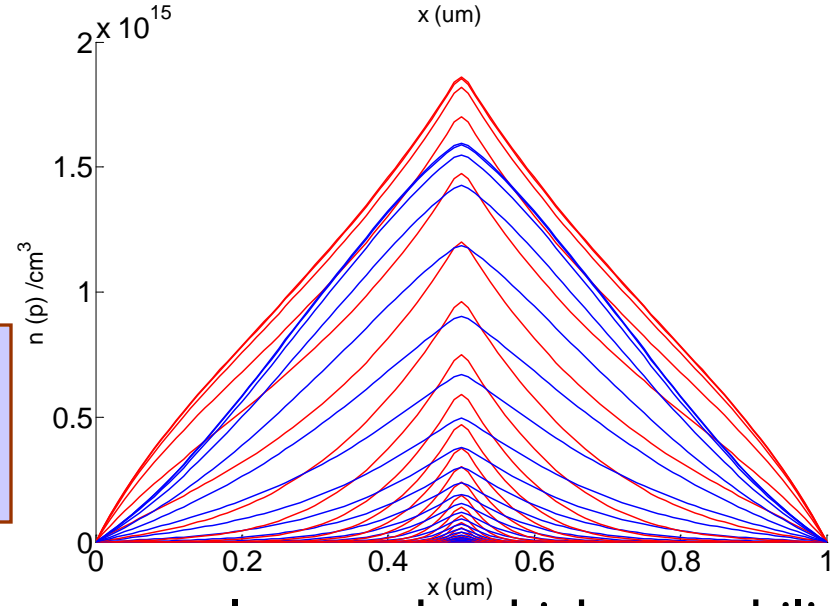
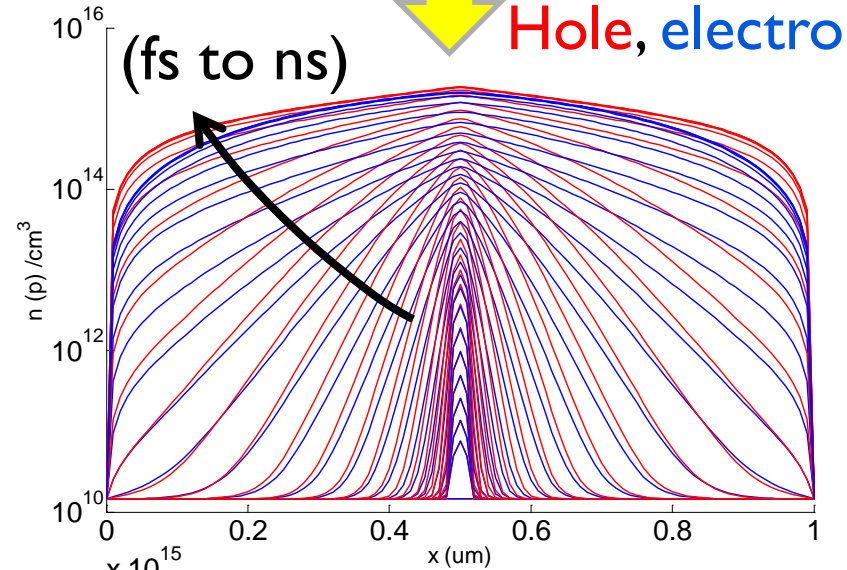
Assume that the incident flux is $1.6 \text{ fC/micron} = 10^{15} / \text{sec/micron}$,
Or $10^{27} \text{ photon /cm}^3 / \text{sec}$ is the incident density
Flux = $10^{27} \times 0.1 \times 10^{-4} = 10^{22} \text{ photon /cm}^2 / \text{sec}$
= $6.4 \text{ nC/cm}^2 / \text{sec}$ (same number of e-h pair)
 $n_i = 10^{10} / \text{cm}^3$, no recombination

Localized photo generated e-h pairs



$$-\frac{d^2V}{dx^2} = \frac{dE}{dx} = \frac{q}{\epsilon} (p - n + N_D - N_A)$$

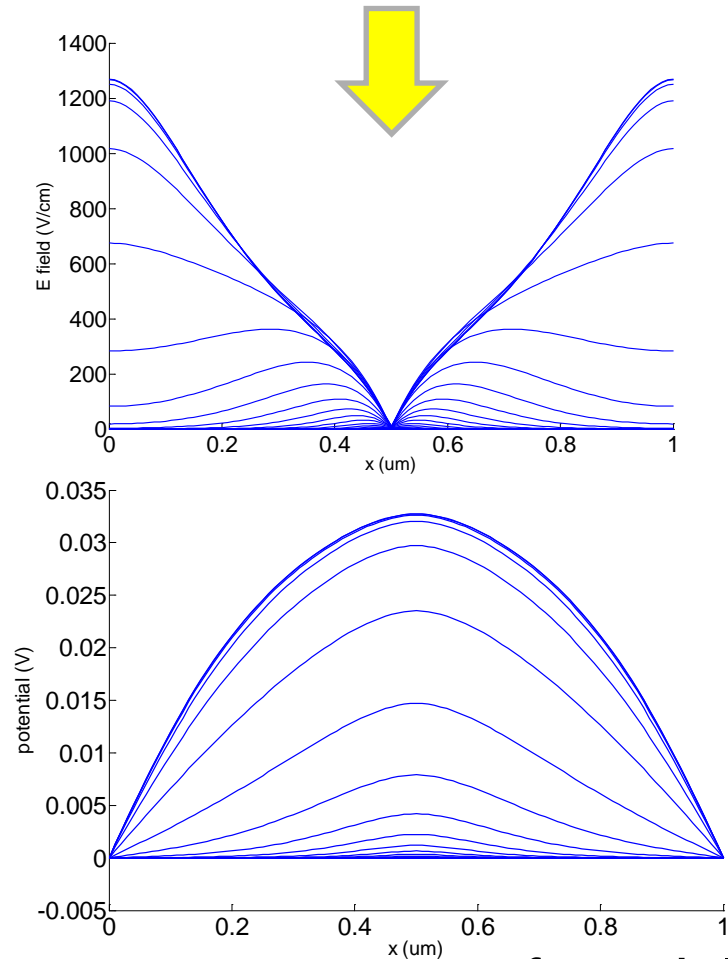
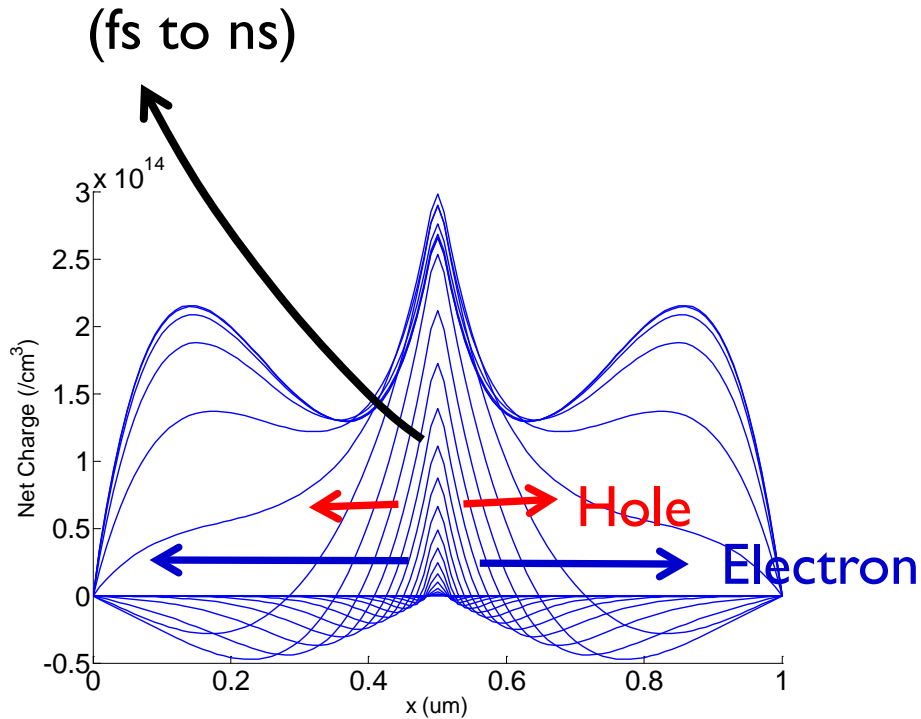
↓ Hole, electron



At steady state, more holes than electron, as electron has higher mobility

Importance of Electrostatics

$$-\frac{d^2V}{dx^2} = \frac{dE}{dx} = \frac{q}{\epsilon} (p - n + N_D - N_A)$$



Dynamic charge separation in an importance source of instability ...

Outline

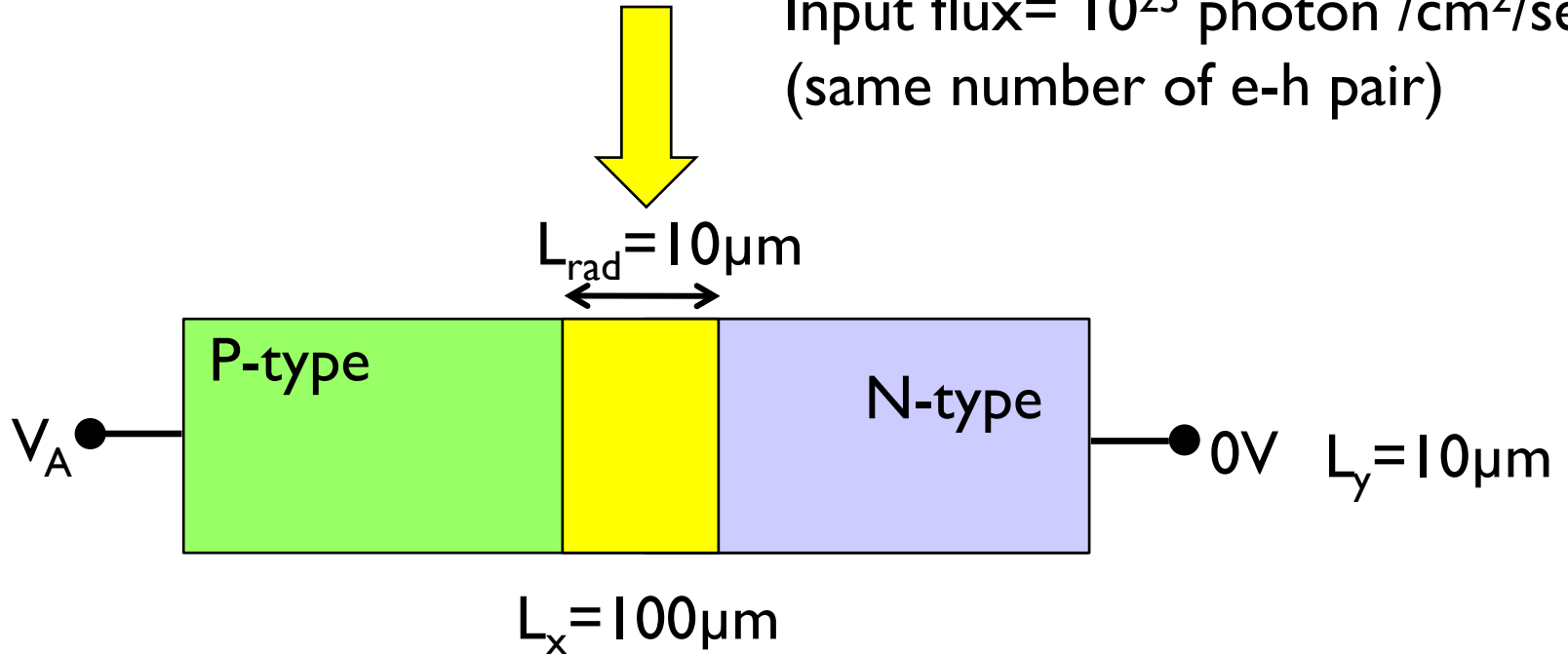
1. Sources of radiation
2. Charge generation and potential fluctuation
3. Junctions and critical charges
4. Conclusions

For p-n junction: even with very low lifetime

$$G = 10^{28} \text{ photon /cm}^3/\text{sec}$$

$$\text{Input flux} = 10^{25} \text{ photon /cm}^2/\text{sec}$$

(same number of e-h pair)

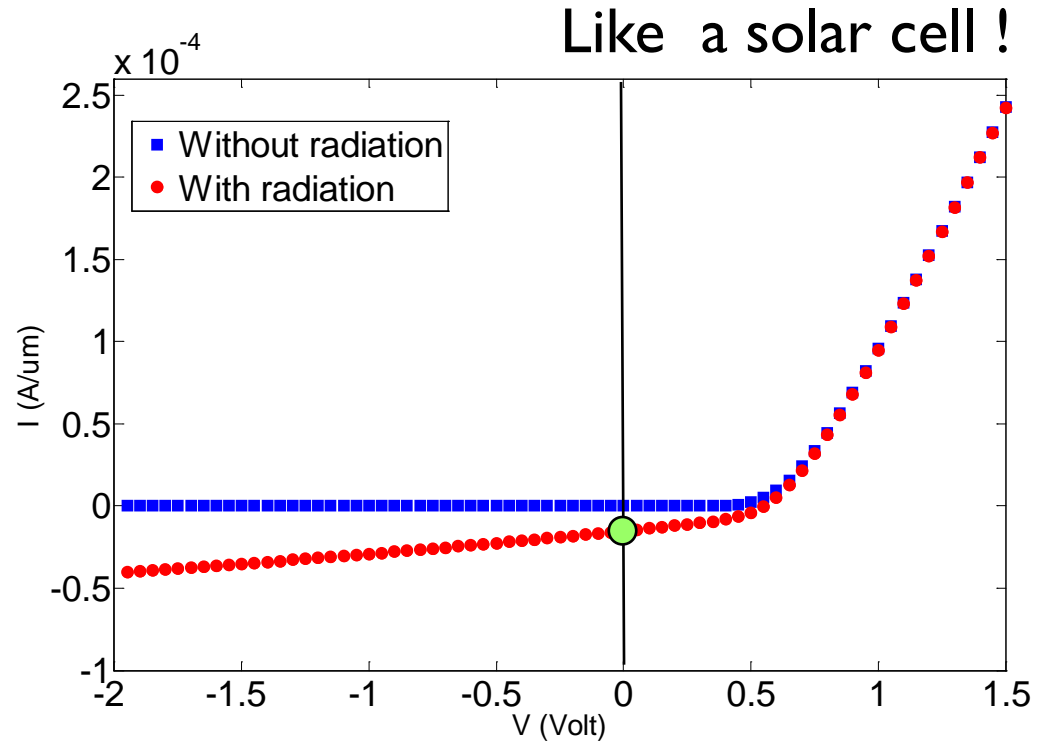
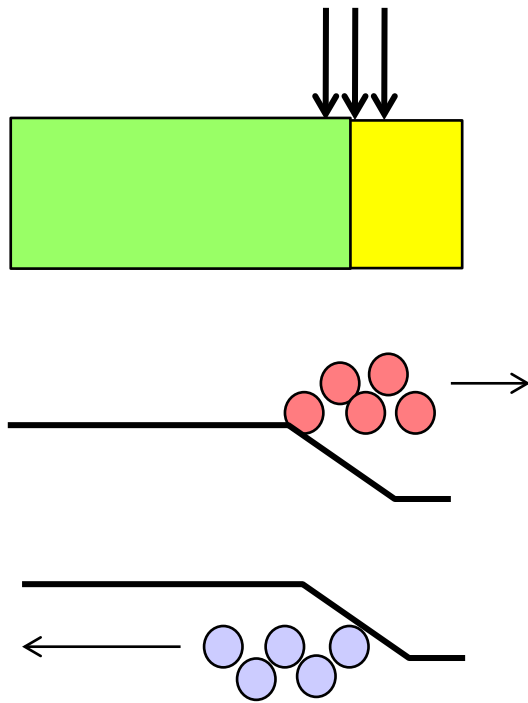


$$N_A = N_D = 10^{18} /\text{cm}^3$$

Minority carrier lifetime, $\tau = 10^{-12} \text{ sec}$

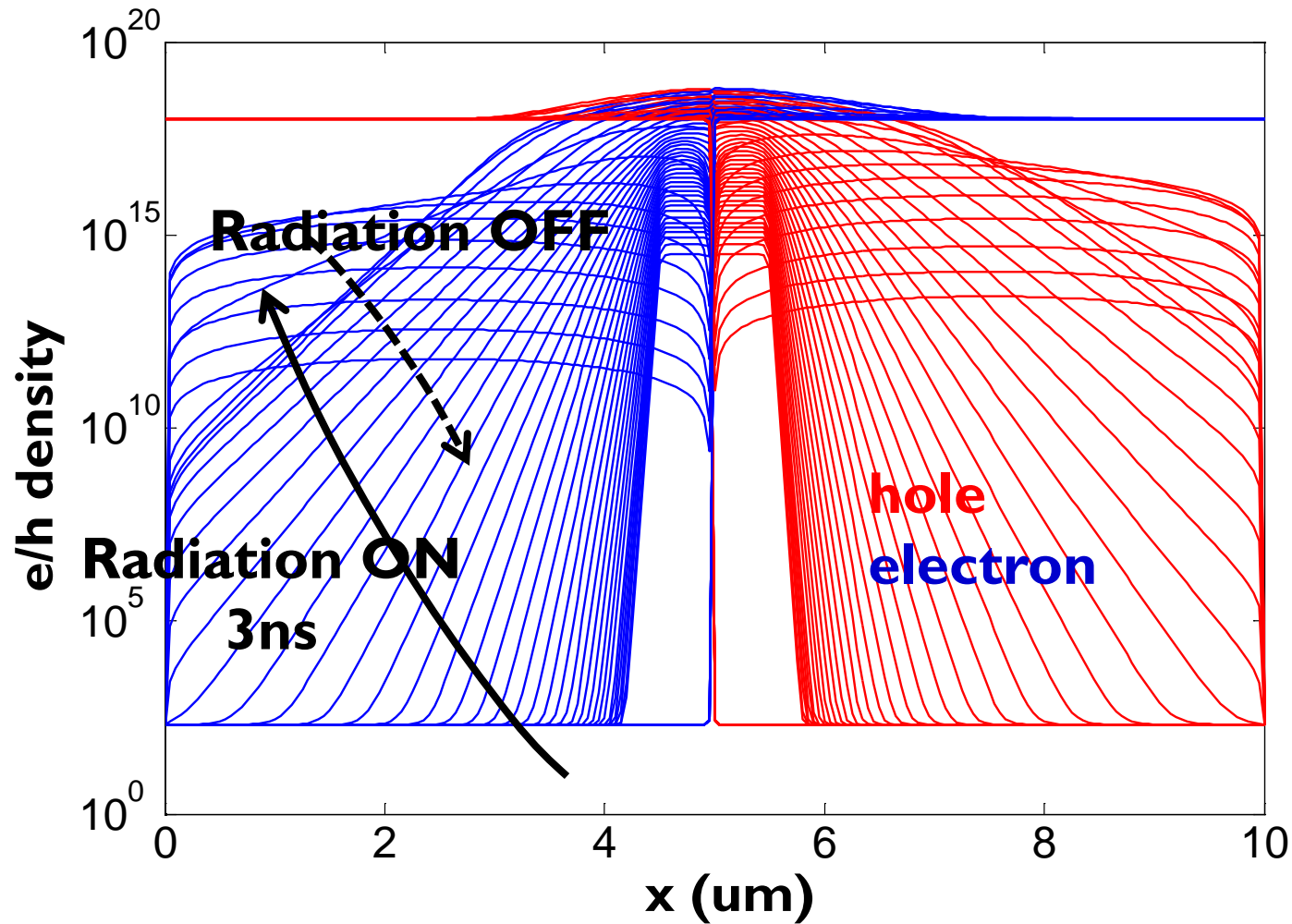
$$\Rightarrow L_D \text{ (estimated)} = \sqrt{(6.5 \text{ cm}^2/\text{sec} * 10^{-12} \text{ sec})} \sim 0.025 \mu\text{m}$$

Photo generated e-h pairs at Junctions

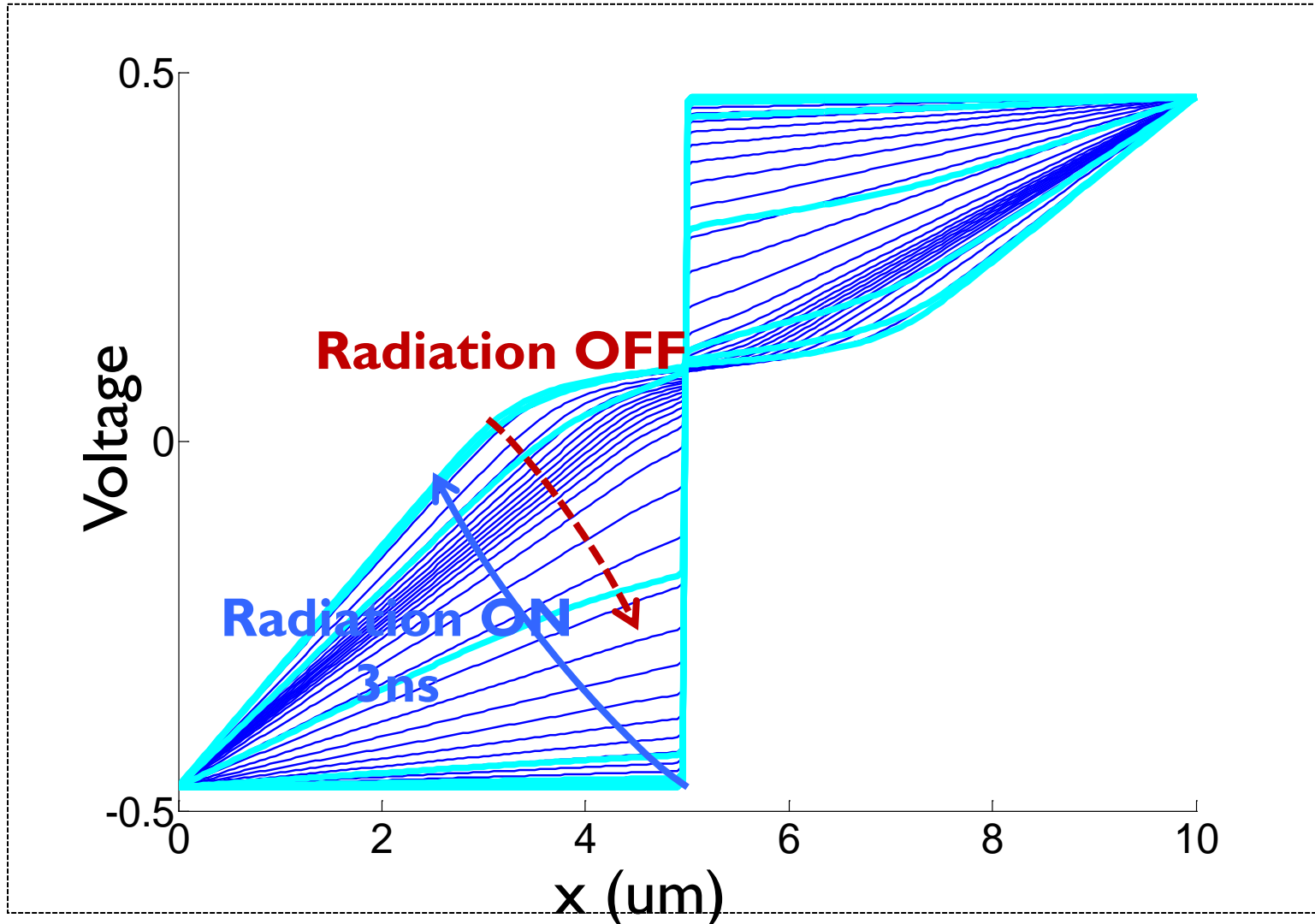


$$-\frac{d^2V}{dx^2} = \frac{dE}{dx} = \frac{q}{\epsilon} (p - n + N_D - N_A)$$

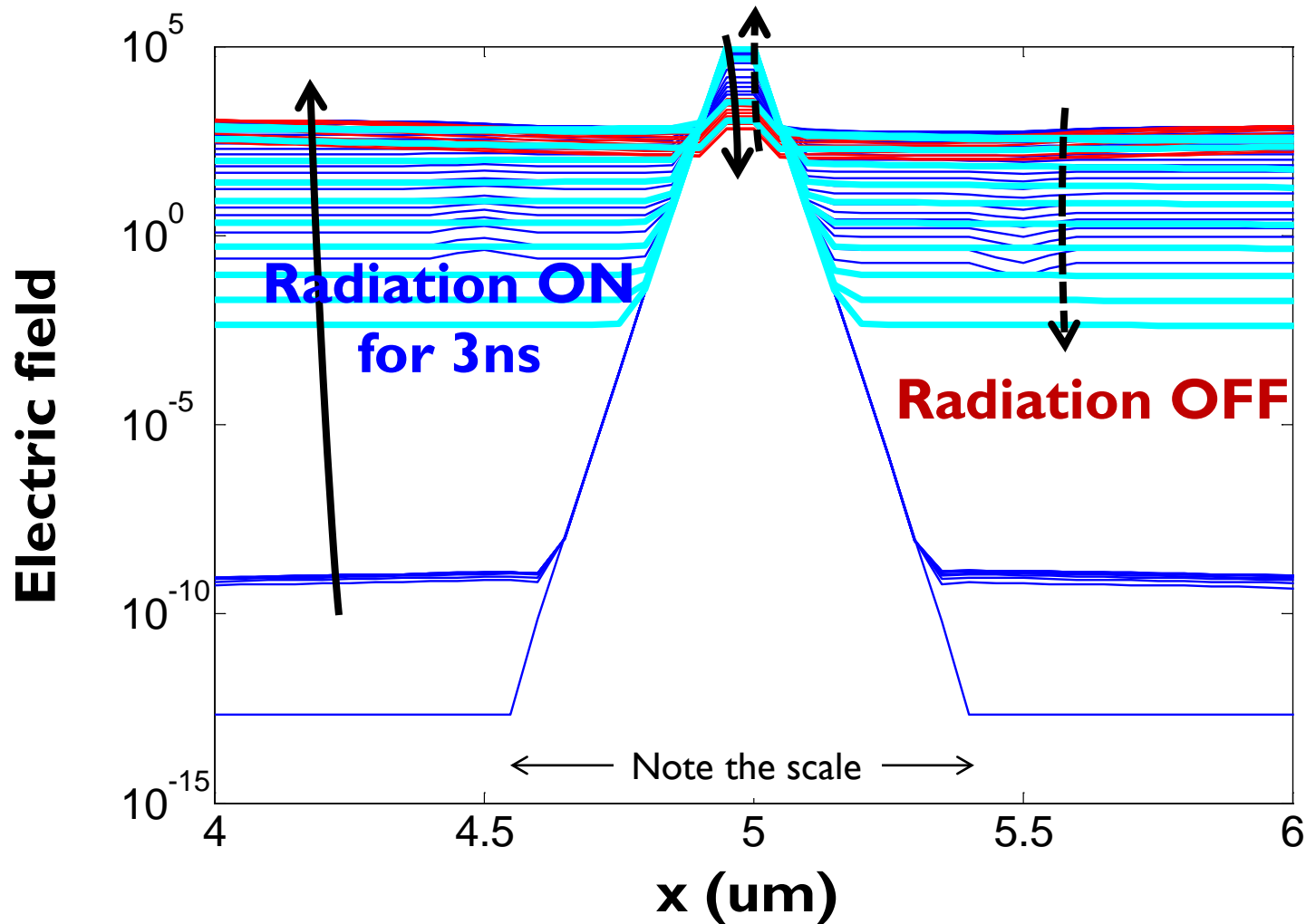
Electron/hole densities



Potential before and after radiation



E- field before and after radiation



Critical Charge: Qcrit

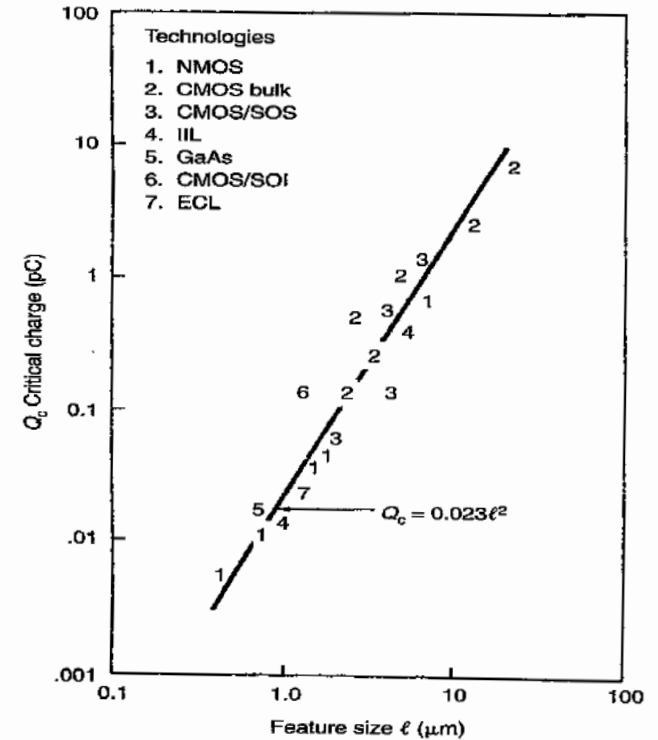
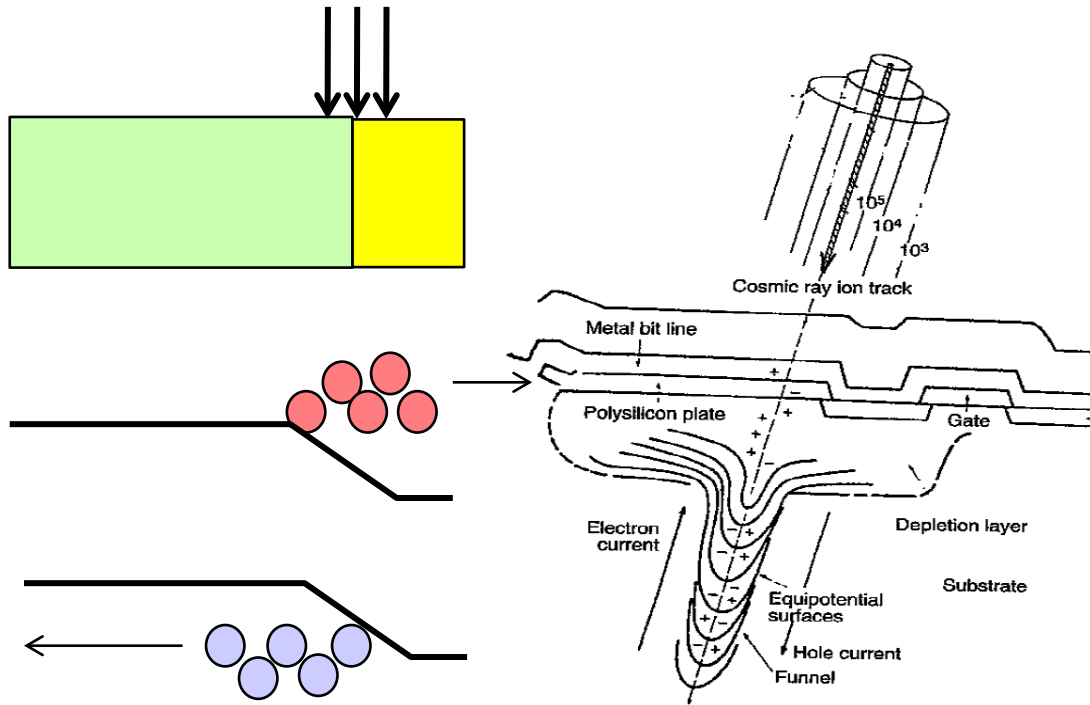


Fig. 7-15 Critical charge for SEU (Q_c) vs feature size in different devices. From E. L. Peterson and P. Marshall, *J. Rad. Eff. Res. and Eng.* (Jan 1989).

$$-\frac{d^2V}{dx^2} = \frac{dE}{dx} = \frac{q}{\epsilon} (p - n + N_D - N_A)$$

Critical charge in devices

Circuit design and modeling for soft errors

A. KleinOsowski
E. H. Cannon
P. Oldiges
L. Wissel

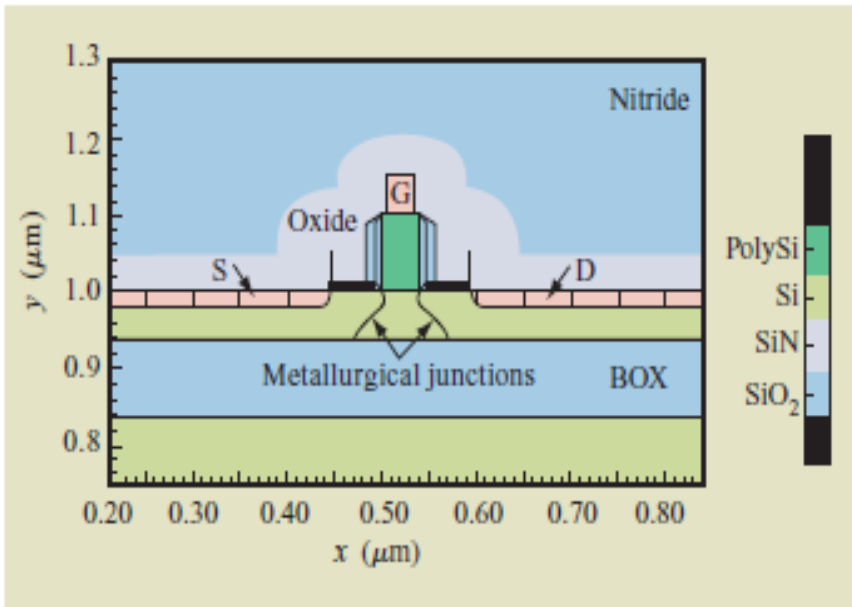


Figure 1

Modeled structure of a 65-nm SOI technology n-FET. (BOX: buried oxide.)

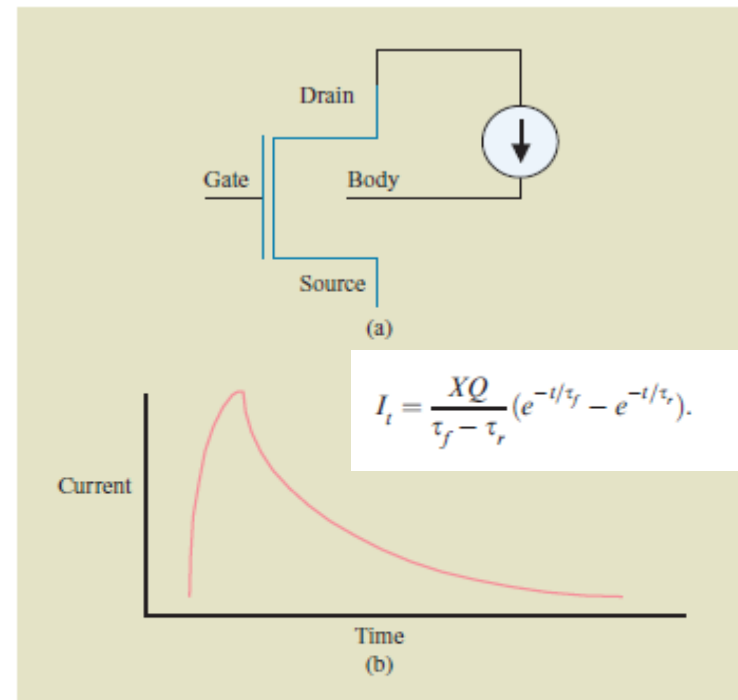


Figure 3

(a) Method for connecting injected current pulse into circuit simulation for SOI Q_{crit} modeling. (b) Current pulse wave shape used for circuit modeling.

Critical Charge in circuits

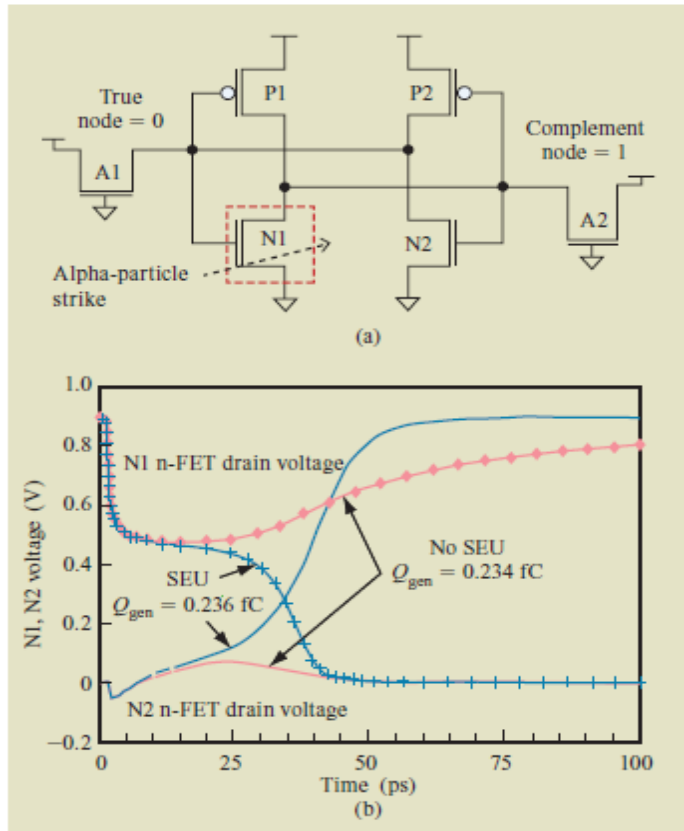


Figure 2

(a) Typical six-transistor SRAM circuit. (b) Evolution of the node voltages for an alpha-particle strike in an off-state n-FET pull-down device (N1) of a 65-nm SOI SRAM cell at $V_{dd} = 0.9$ V. (SEU: single-event upset.)

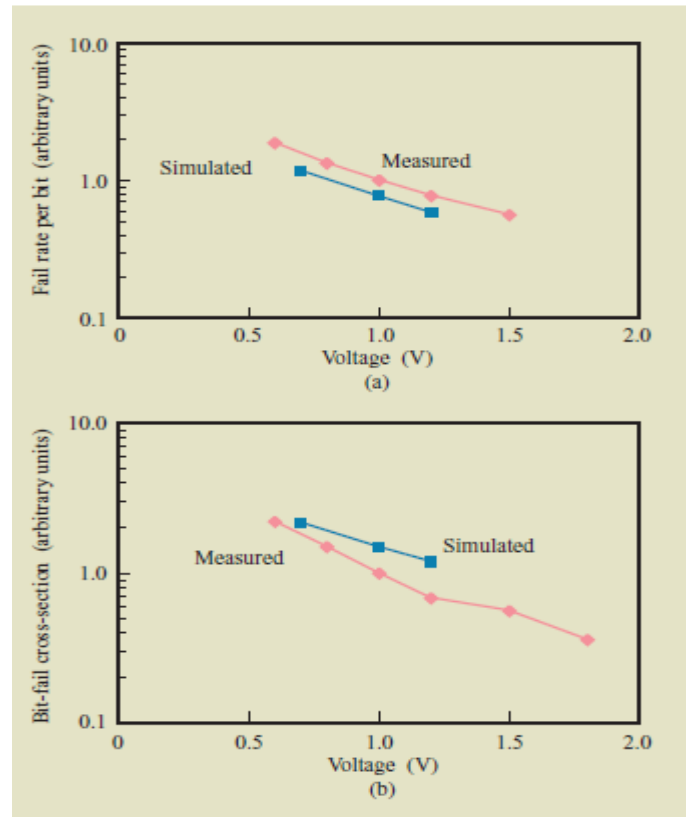


Figure 6

(a) Alpha-particle sensitivity for 90-nm bulk technology SRAM: measurements with a Th-232 foil and Monte Carlo simulations. (b) 150-MeV proton sensitivity for 90-nm bulk technology SRAM: measurements in a proton beam and Monte Carlo simulations. (Reproduced from [8].)

Conclusions

1. Soft error is related to generation of electron-hole pairs that perturbs the electrostatics and temporarily destroys the logic state of the devices.
2. Soft error is reversible: memory can be rewritten; the logic state can be recomputed.
3. The asymmetry of electron hole transport (either due to difference in mobility, doping, or junction) that causes soft errors.
4. Soft errors can be mitigated by clever device design as well as redundant circuit approaches. The solutions are costly – therefore DARPA and NASA uses special radiation hardened components.
5. The story of Icarus and the role radiation damage!



References

R.C. Bauman wrote an wonderful review on packing related soft-errors, see Soft errors in Advanced Semiconductor Devices, TDMR, 1, 1, 2001.

IBM Journal of Research and Development has published two special issues on radiation induced damage. They contain wonderful sources of information.

The following book gives a very good introductory overview of the physics of soft errors” SER – History, Trends, and Challenges, A guide for designing with memory lcs, J. F. Ziegler and H. Puchner, Cypress. 2004.

A number of figures are taken from various sources:

<http://www.meyerinst.com/html/oem/in-focus/default.htm> (package)

Lfm.mit.edu (nuclear explosion)

<http://www.military-heat.com/tag/jet/page/2/> (Fighter jet)

www.wired.com (Sun)

Iracus picture taken from

http://wings.avkids.com/Book/Myth/Images/icarus_03.jpg

Review Questions

1. How does Earth's magnetic field deflect charged particles? What are its consequences for geography-specific radiation damage?
2. What is a Pfofzer curve. How is it used?
3. What does the isotropy of radiation say about the origin of galactic radiation sources?
4. A special type of high-energy particle was used to validate Einstein's theory of relativity. Are those particles a radiation concern for ICs?
5. Nuclear reaction from *packaging* leads to a large number of upsets. Why?
6. Can a soft-error occur in an intrinsic CNT? Why or why not?
7. Define "critical charge". Is it a property of a device or a property of the radiation source?